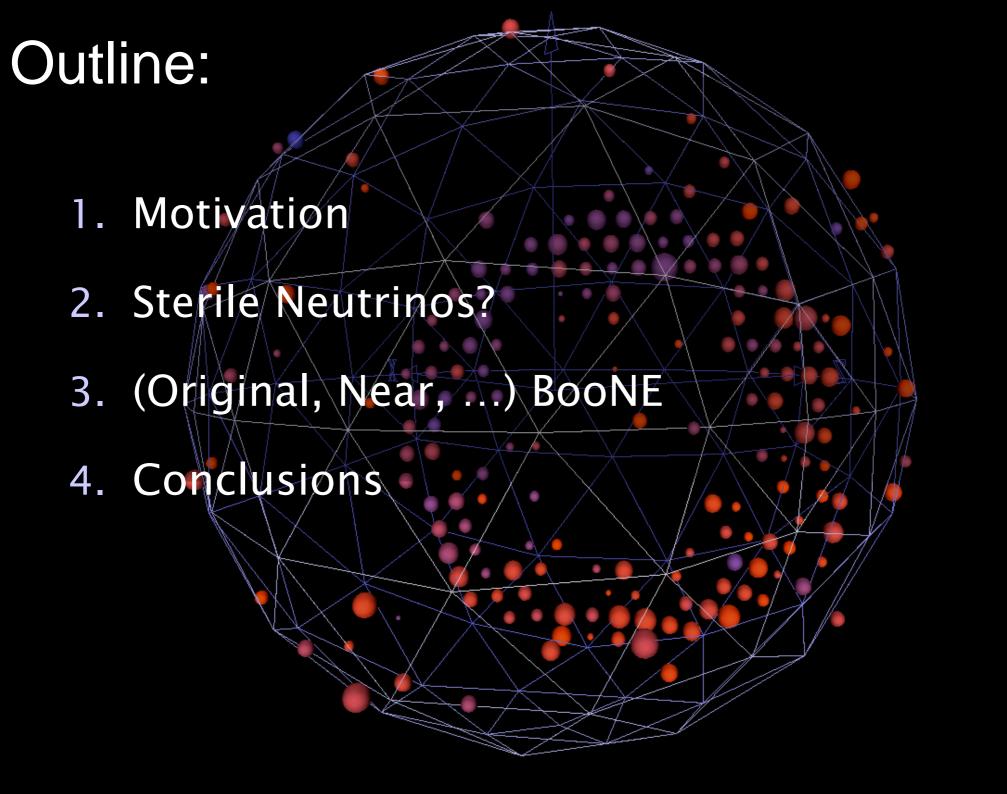


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P-25 Subatomic Physics Group

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#### Neutrino Oscillations

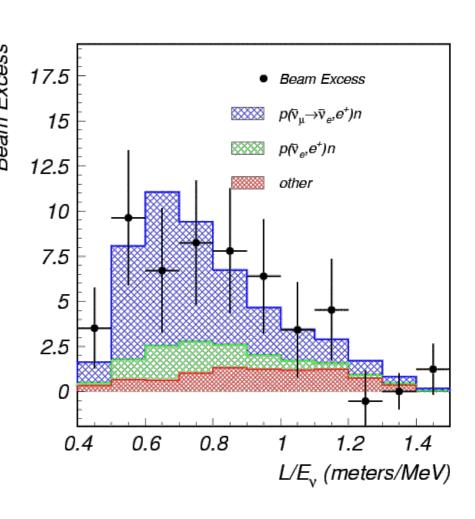
- The oscillation patterns between the 3 known active neutrino species have been demonstrated by a number of experiments over the last two decades:
  - SNO, Kamland
  - Super-K, K2K, MINOS
- Armed with that knowledge, measurements of neutrino behavior outside the standard 3 generations of active neutrinos indicate new physics:
  - LSND indicates that new physics may be operating
- Interpretations of such a non-standard result probe some deep theoretical issues, for example:
  - Light sterile neutrinos, neutrino decays, CP and/or CPT violation, Lorentz invariance, Extra dimensions

The investigation of neutrino oscillations at the <1% level is unique in its physics reach

### Excess Events from LSND

- LSND found an excess of  $\overline{\nu}_e$  in  $\overline{\nu}_\mu$  beam
- Signature: Cerenkov light from e+ with delayed n-capture (2.2 MeV)
- Excess:  $87.9 \pm 22.4 \pm 6.0$  (3.8 $\sigma$ )
- The data was analysed under a two neutrino mixing hypothesis\*

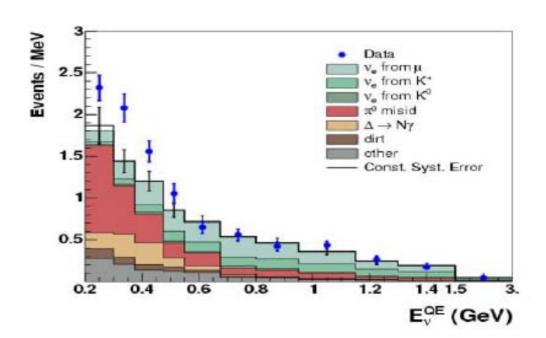
$$P(\overline{\nu}_{\mu} \to \overline{\nu}_{e}) = \sin^{2}(2\theta) \sin^{2}\left(\frac{1.27 L \Delta m^{2}}{E}\right)$$
$$= 0.245 \pm 0.067 \pm 0.045 \%$$



\*at least 5 neutrinos are required to accommodate all experiments

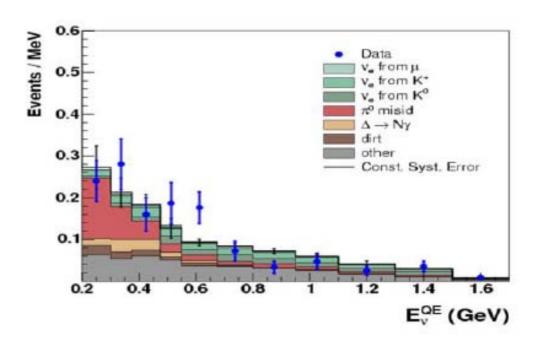
### MiniBooNE Results

### Neutrino Mode

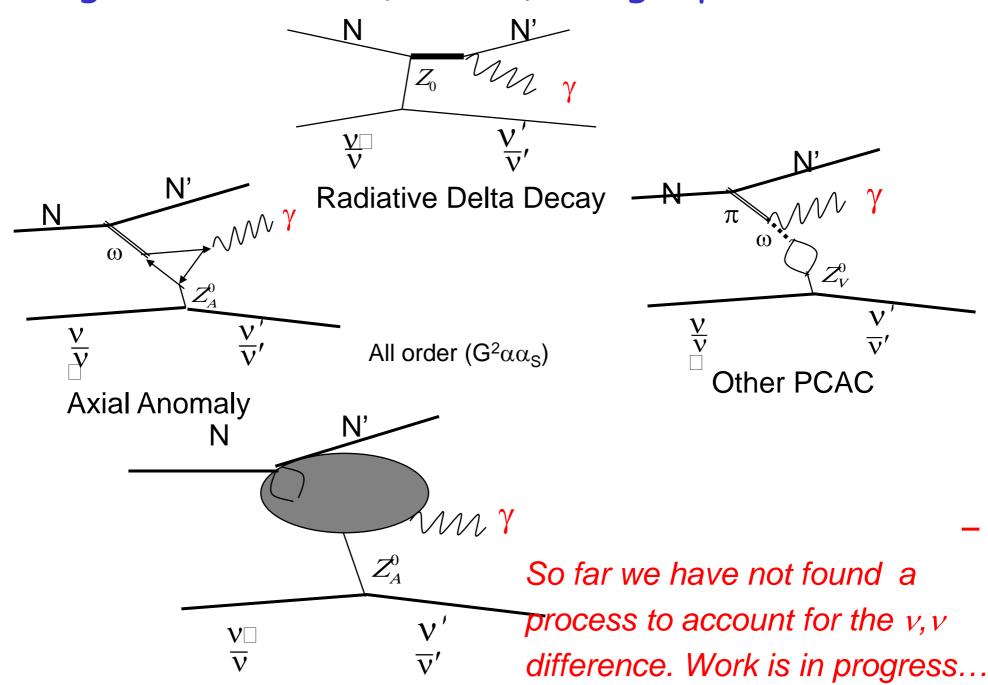


### Anti-neutrino Mode

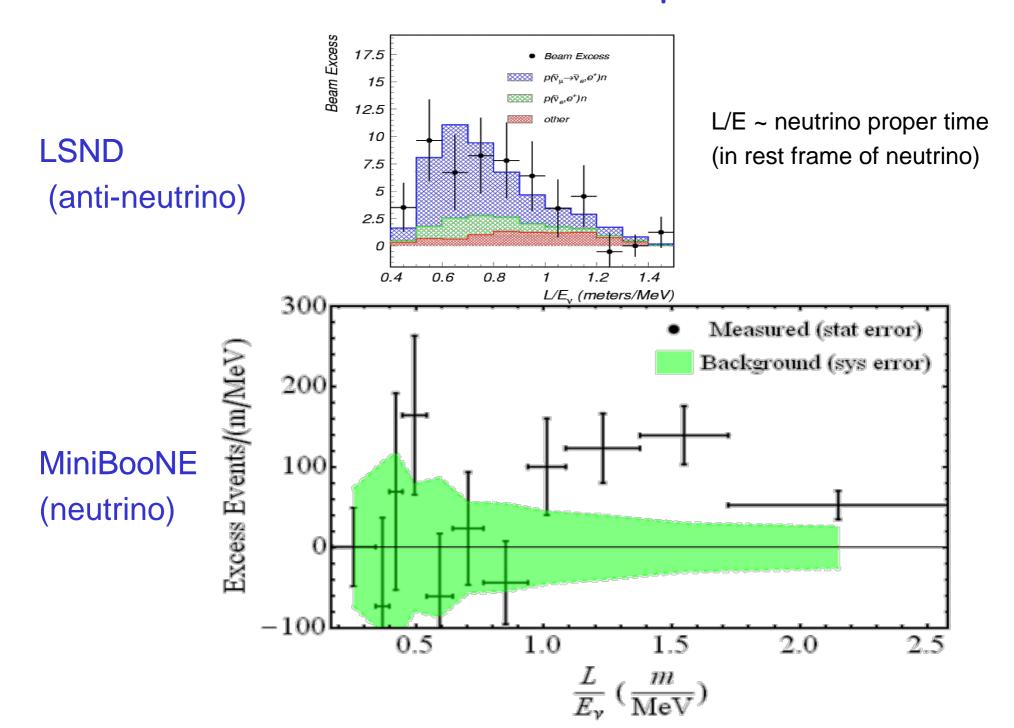
(Not yet sensitive to LSND)



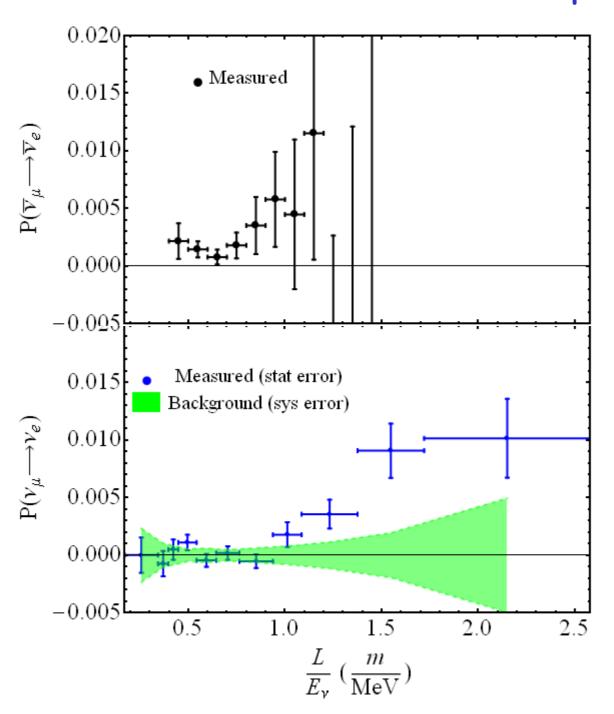
# Backgrounds: Order( $\alpha \times NC$ ), single photon FS



## MiniBooNE and LSND Comparison



## LSND and MiniBooNE oscillation probabilities



## LSND and MiniBooNE oscillation probabilities

My own attempts to reconcile Data:

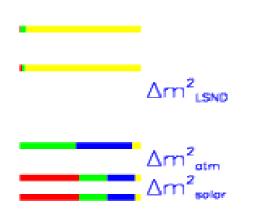
#### "high-low" solution

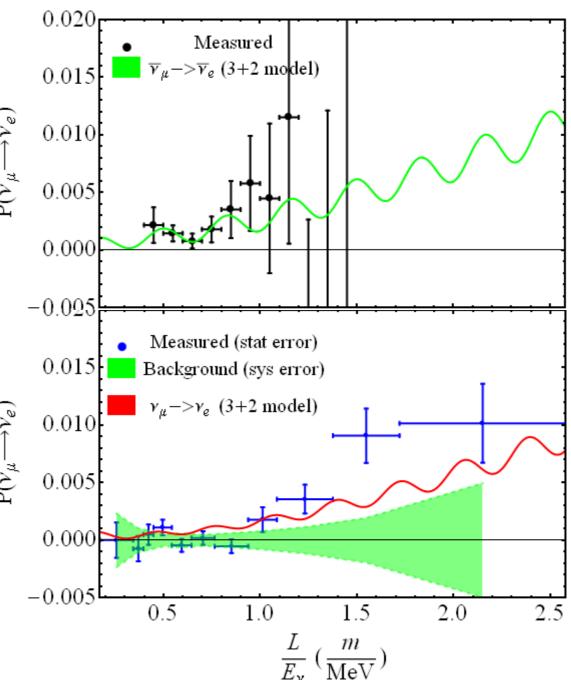
3+2 model (suggestive)

$$\Delta m_a^2 = 7.5 \text{ eV}^2, P_a = 0.015$$

$$\Delta m_b^2 = 0.25 \text{ eV}^2, P_b = 0.065$$

$$\phi_{CP} = 1.3 \, \mathrm{rad}$$





LSND and MiniBooNE oscillation probabilities

My own attempts to reconcile Data:

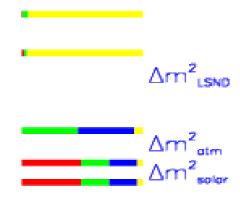
#### "low-low" solution

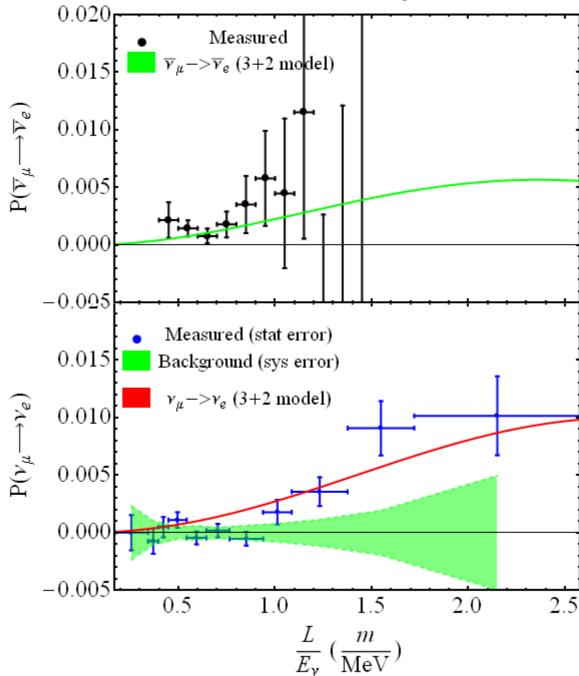
3+2 model (suggestive)

$$\Delta m_a^2 = 0.5 \text{ eV}^2, P_a = 0.04$$

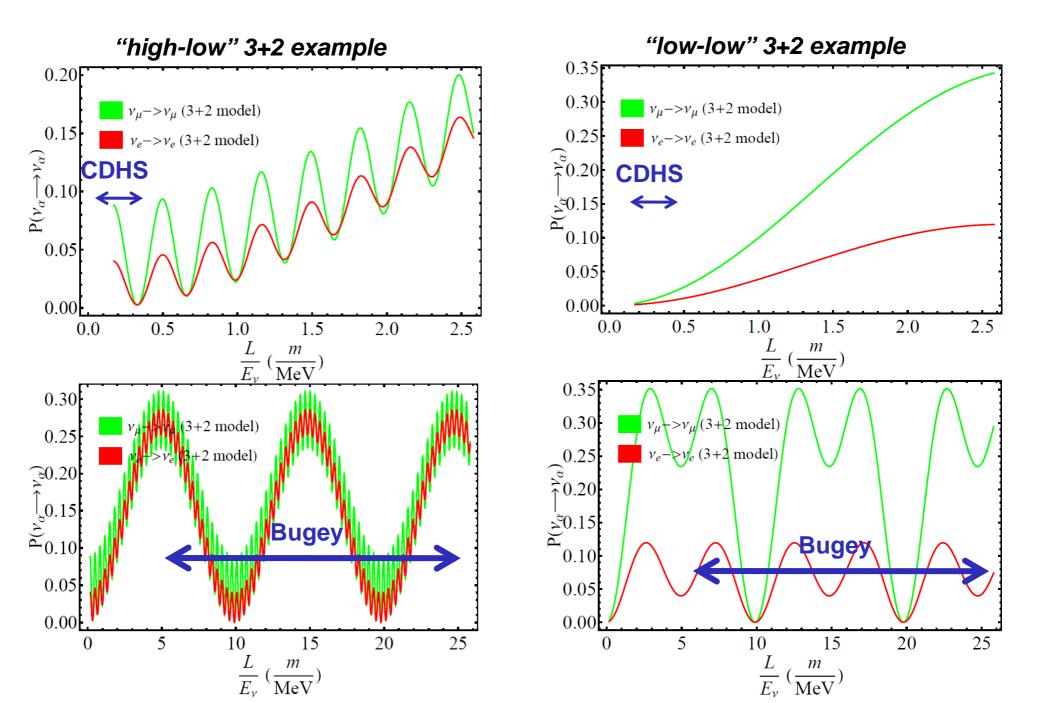
$$\Delta m_b^2 = 0.25 \text{ eV}^2, P_b = 0.025$$

$$\phi_{CP} = \frac{\pi}{2} \text{ rad}$$



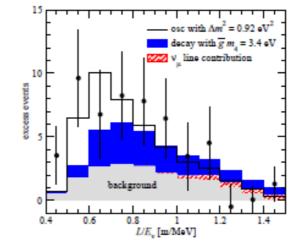


# Disappearance oscillation probabilities

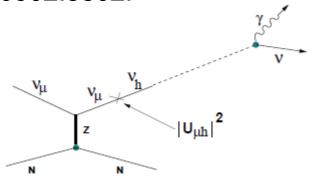


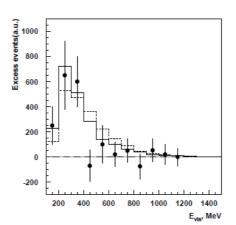
## Neutrino Decays?

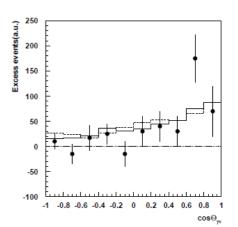
- With the addition of heavy sterile neutrinos the effects of sterile neutrino decays become important
- The combined effect has been shown to separately accommodate LSND and MiniBooNE results
  - arXiv:hep-ph/050521:



– arXiv:0902.3802:







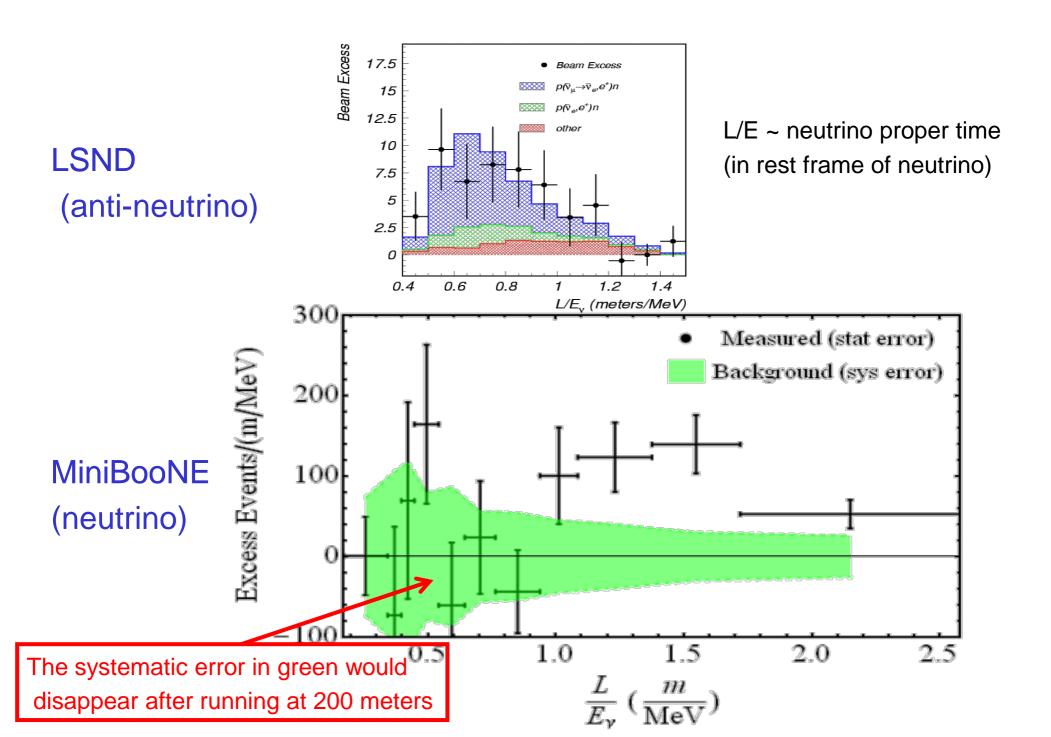
## Resolving the MiniBooNE Low Energy Excess

- ➤ Moving the MiniBooNE detector to 200m (~4M\$) (or building a new detector at 200m (~\$8M\$))
  - ➤ Accumulate a sufficient data sample in < 1 year
  - ➤ will dramatically reduce systematic errors (low energy excess is ~ 6 sigma significance with statistical errors only.
  - ➤ Can study L dependence of excess: backgrounds scale as 1/L\*\*2, oscillation signal as sin²(L/E), and decay as L/E.
- ➤ MicroBooNE (next talk...):
  - is a 70 ton liquid argon time projection chamber planned for the booster neutrino beam line
  - can differentiate single gamma-rays from electrons (MiniBooNE cannot do this)

#### Near BooNE Detector

- In the low-e analysis we are dominated by systematic errors
  - The low-e excess is ~ 6 sigma statistically, but only ~3 sigma including systematic errors
    - it might be due to a unforeseen background or...
    - it might be due to new, non-standard physics
- Running MiniBooNE for 1 year at a position 200 meters would increase the data rate per pot x6 and:
  - Make flux, cross section, and optical model systematic errors small in the 200 meter/500 meter comparison
  - → Demonstrate at ~ 5 sigma level whether or not the low-e excess depends on 1/L² or (L,E) e.g. oscillations

### MiniBooNE at 200 meters

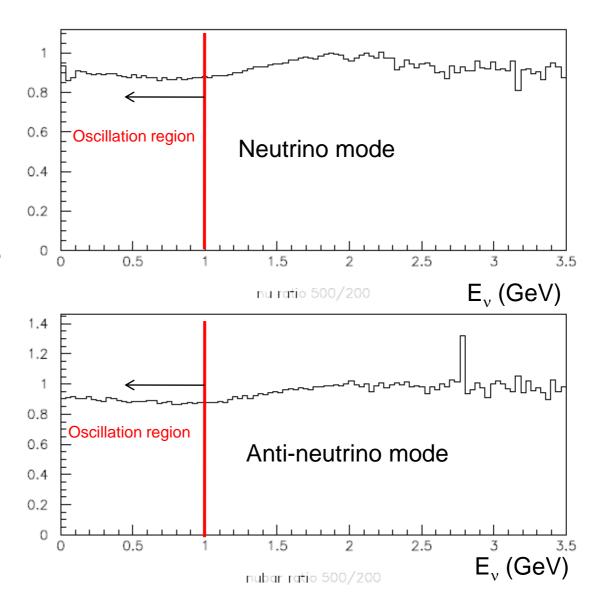


### Systematic Errors:

- Neutrino Cross Sections
  - target material is the same (mineral oil)
- Detector Efficiencies
  - With the same detector, we should have nearly identical detection efficiencies
- Flux ?
  - the opening angles of neutrinos from pion or kaon decays are tens of mrad or greater
  - Most pions and kaons contribute to both the 200 and 500 meter fluxes

### Far to Near Neutrino Flux Ratios at 200 m

MiniBooNE Far/Near fluxes Scaled by 1/r<sup>2</sup>



#### Statistical Errors

- In order to achieve reasonable sensitivity ~ ½-1 year of running would be required for each focus
  - Current proton delivery rates of 2 x 10<sup>16</sup> protons/hour give ~1.75x10<sup>20</sup> protons/year
  - Current 6.5 σ statistical significance translates to ~5 σ statistical significance for one year of running in nu mode

### Options for Near BooNE Detector

- Transport existing MiniBooNE detector (~80 tons) to new location 150-200 meters from BNB target (~4M\$)
- Dismantle existing MiniBooNE detector and construct a new detector at 150-200 meters. (~4M\$)
- Construct brand new detector at 150-200 meters (~8M\$)

#### Lift of 260 ton Generator





Transporting 550 ton Coker Drum from ship to crane hook

#### Conclusion

- Moving MiniBooNE to 200 meters and running for one year would resolve whether or not the low-e excess is due to a (L,E) dependent phenomena at the ~ 5 sigma level
- It would also provide a high statistics, low systematic error numu/numu-bar disappearance measurement
- The timing of the project is ideal for post-antineutrino running in the BNB
- MicroBooNE will at the same time run to look for excess gamma events
- A "LSND"-like detector at the SNS (OscSNS) would directly test the LSND excess